CS 563 Advanced Topics in Computer Graphics *Volume Scattering*

by Paulo Gonçalves de Barros

Introduction

- Scenes in vacuum
- Real-life
 - Atmosphere
 - Smoke
 - Haze
 - Clouds

Introduction

- Scenes in vacuum
- Real-life
 - Atmosphere
 - Smoke
 - Haze
 - Clouds
- Volume scattering
 - Participating media
 - Its effect on light rays passing through it

Summary

- Volume Scattering processes
- Phase Functions
- Volume Interface
 - Homogeneous Media
 - Varying Density Volumes
 - Volume Aggregates

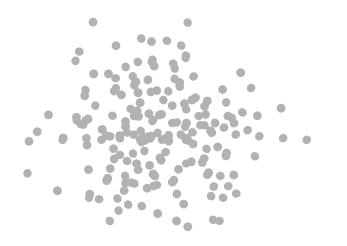
- Pariticipating media properties
 - Absorption
 - Emission
 - Scattering
 - Out-scattering
 - In-scattering
 - Homogeneous or Inhomogeneous

Absorption

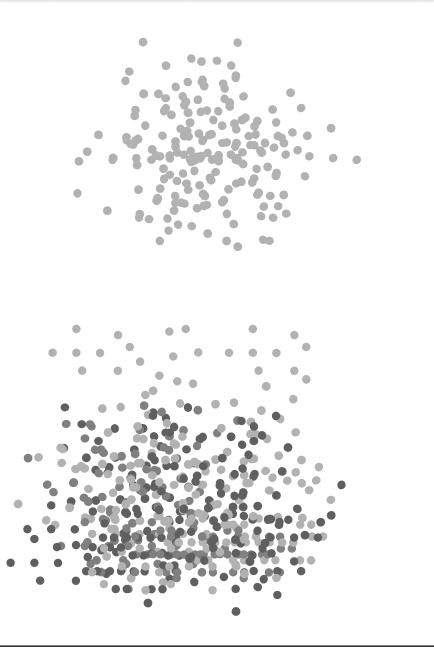
Volume Scattering Processes Absorption Emission

Volume Scattering Processes Scattering

- Homogeneous
 - Constant particle density
 - Uniform particle types distribution



- Homogeneous
 - Constant particle density
 - Uniform particle types distribution
- Inhomogeneous
 - Varying particle density
 - Varying particle distribution



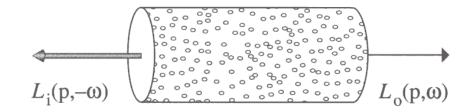
Absorption

- Light is absorbed by medium
- Ray radiance decreases through the medium

- Absorption crossed section s_a
 - Light absorption probability density per unit distance traveled in medium
 - Units ? m⁻¹
 - dt ? through-medium-travel unit
 - Values may be larger than 1
 - Influence factors
 - Position (p)
 - Direction (?)
 - Spectrum

- Change in radiance per unit
 - Difference between incoming and outgoing radiance

$$dL_o(p, \mathbf{w}) = L_o(p, \mathbf{w}) - L_i(p, \mathbf{w})$$



- Change in radiance per unit
 - Difference between incoming and outgoing radiance

$$dL_o(p, \mathbf{w}) = L_o(p, \mathbf{w}) - L_i(p, \mathbf{w})$$

Negative fraction of L_i

$$dL_o(p, \mathbf{w}) = -\mathbf{S}_a(p, \mathbf{w})L_i(p, -\mathbf{w})dt$$

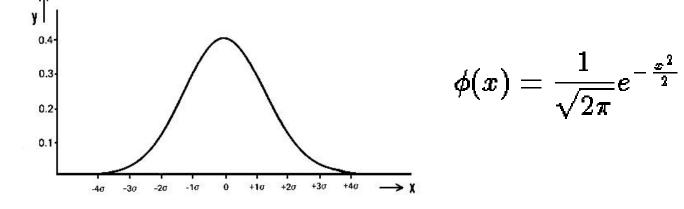
$$L_{i}(p,-\omega)$$

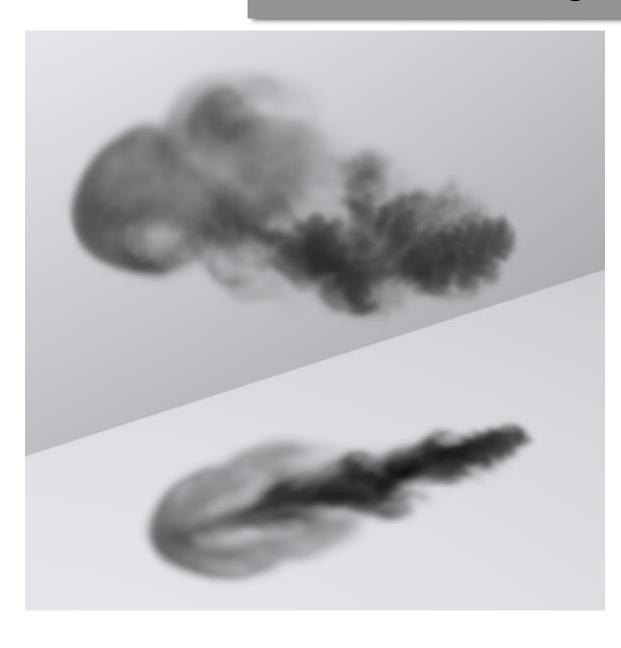
- Absorbed radiance
 - Traveled a distance d through medium

s_a integrated along d $L(\mathbf{p} + \mathbf{w}d, \mathbf{w}) = L(\mathbf{p}, \mathbf{w})e^{-\int_0^d \mathbf{s}_a(\mathbf{p} + \mathbf{w}t, \mathbf{w})dt}$ Probability density function

- Absorbed radiance
 - Traveled a distance d through medium

Normal probability density function (Gaussian)





Emission

- Light is emitted by the medium
- Emitted radiance: $L_{ve}(p, w)$
 - Independent of incoming light

Emission

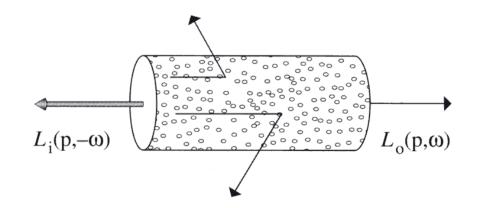
- Light is emitted by the medium
- Emitted radiance: $L_{ve}(p, \mathbf{W})$
 - Independent of incoming light
- Change in radiance per unit $dL_o(p, \mathbf{w}) = L_{ve}(p, \mathbf{w})dt$



Out-scattering

- Light is scattered out of the path of the ray
- Probability density for scattering: s_s
- Reduction in radiance is given by

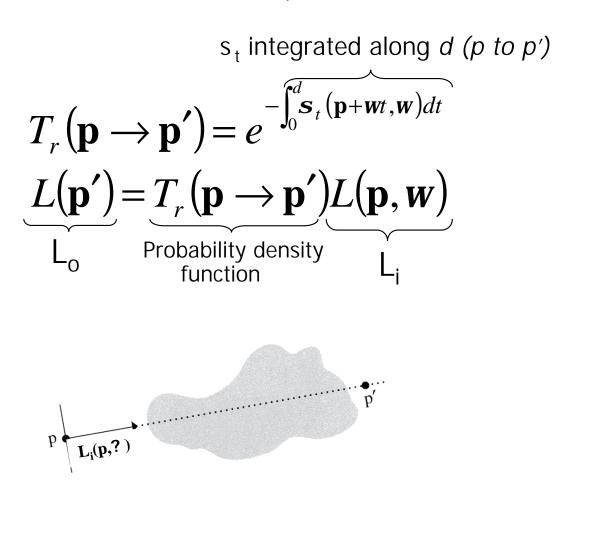
 $dL_o(p, \mathbf{w}) = -\mathbf{S}_s(p, \mathbf{w})L_i(p, -\mathbf{w})dt$



- Total radiance reduction
 - Absorption
 - Scattering
- Attenuation or extinction
 - Coefficient: s_t

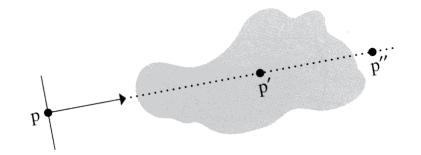
$$\boldsymbol{s}_{t}(\boldsymbol{p}, \boldsymbol{w}) = \boldsymbol{s}_{a}(\boldsymbol{p}, \boldsymbol{w}) + \boldsymbol{s}_{s}(\boldsymbol{p}, \boldsymbol{w})$$

• Change in radiance per unit $dL_o(p, \mathbf{w}) = -\mathbf{S}_t(p, \mathbf{w})L_i(p, -\mathbf{w})dt$ Beam transmittance T_r



Transmittance

- Fraction of light that is transmitted between two points
- Values between 0 and 1
- Properties
 - $Tr(p? \ p) = 1$
 - In vacuum: Tr(p? p') = 1, for all p'
 - Tr(p? p'') = Tr(p? p') Tr(p'? p'')



$$T_r(\mathbf{p} \to \mathbf{p'}) = e^{-\int_0^d \mathbf{s}_t (\mathbf{p} + \mathbf{w}t, \mathbf{w}) dt}$$

Optical thickness

$$t(\mathbf{p} \rightarrow \mathbf{p'}) = \int_0^d \mathbf{s}_t (\mathbf{p} + \mathbf{w}t, \mathbf{w}) dt$$

Homogeneous medium

- s_t is position independent
- Transmittance reduced to Beer's Law

$$T_r(\mathbf{p}\to\mathbf{p'})=e^{-\mathbf{s}_t d}$$

Beer's Law

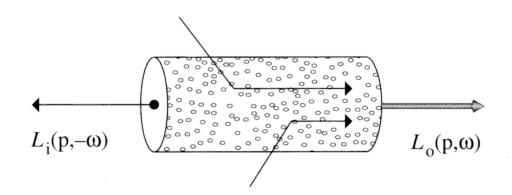
A = alc

In PBRT: $\boldsymbol{s}_t = \boldsymbol{a}(c)$ d = l

- A = amount of light absorbed
- a = Absorption
 coefficient or molar
 absorptivity of medium
- I = distance light travels through medium
- c = Concentration or particle density

In-scattering

- Outside light scatters converging to ray path
- Phase functions to represent scattered radiation in a point



- Phase function (PF)
 - Volumetric analog of BSDF
 - Normalization constraints
 - PF defines a direction's scattering probability distribution

$$\int_{S^2} p(\mathbf{w} \to \mathbf{w'}) d\mathbf{w'} = 1$$

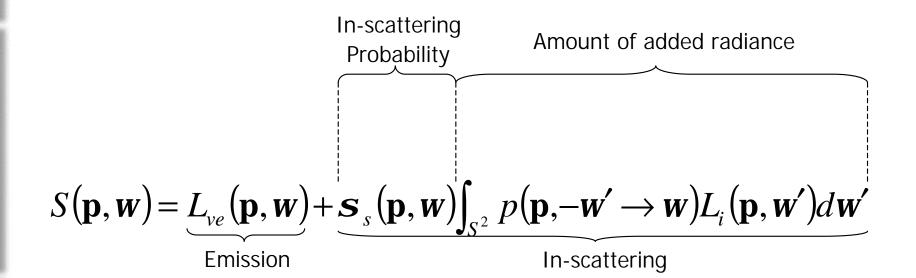
• Change in radiance per unit $dL_{o}(p, \mathbf{w}) = S(p, \mathbf{w})dt$

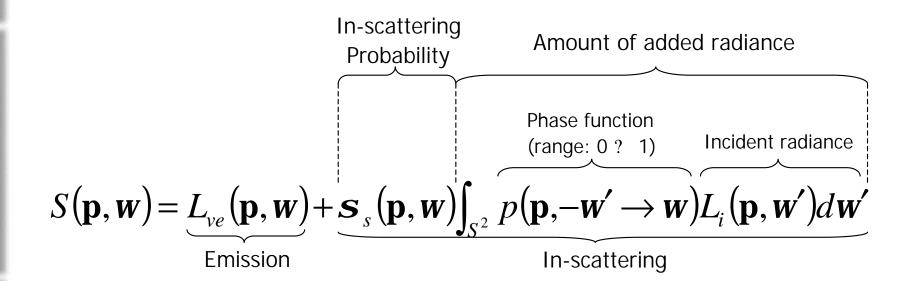
S(p,w) includes volume emission

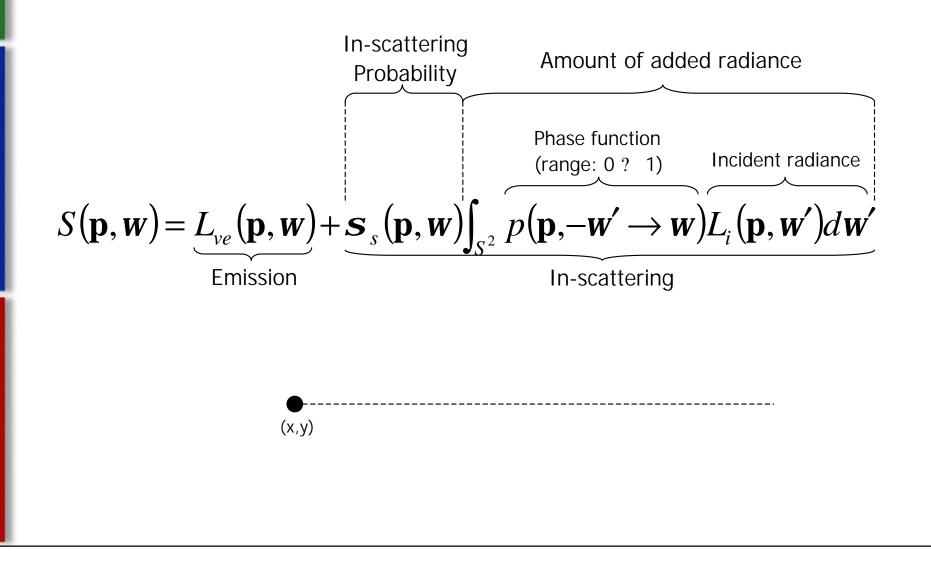
 $S(\mathbf{p}, \mathbf{w}) = L_{ve}(\mathbf{p}, \mathbf{w}) + \mathbf{s}_{s}(\mathbf{p}, \mathbf{w}) \int_{S^{2}} p(\mathbf{p}, -\mathbf{w}' \to \mathbf{w}) L_{i}(\mathbf{p}, \mathbf{w}') d\mathbf{w}'$

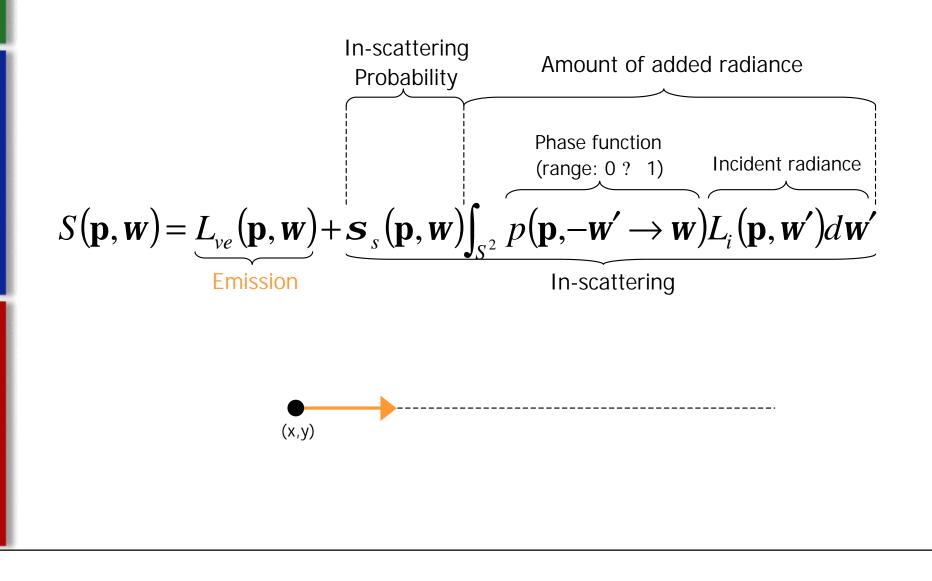
Emission

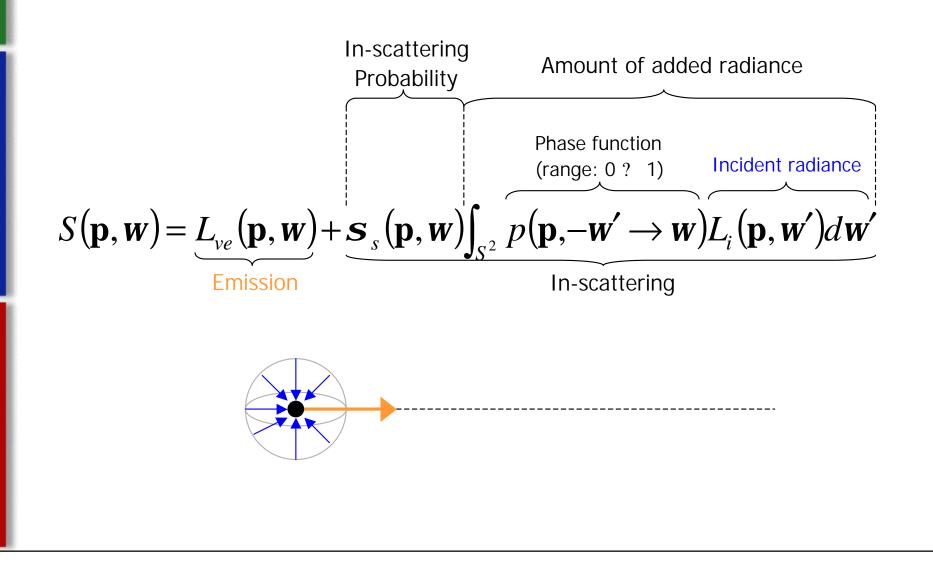
In-scattering

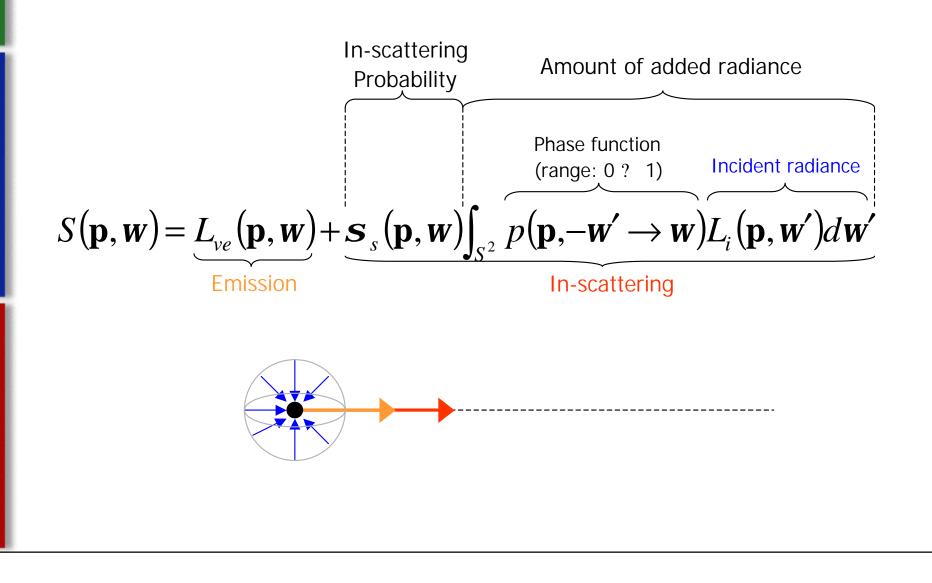






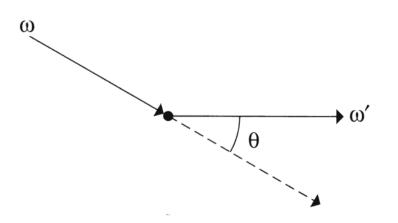




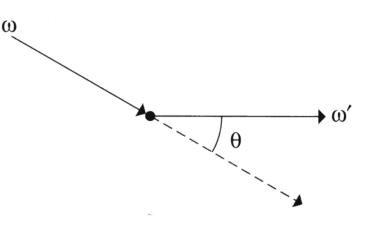


Phase Functions

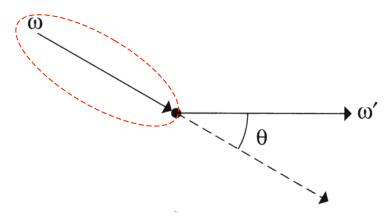
- BSDFs for volume scattering
- Vary complexity according to medium
 - Isotropic
 - Anisotropic



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- Vary complexity according to medium
 - Isotropic
 - Anisotropic
- Properties
 - Direction reciprocity
 - May also be classified as
 - Isotropic uniform scattering
 - Anisotropic variable scattering

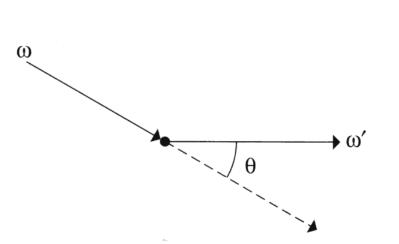


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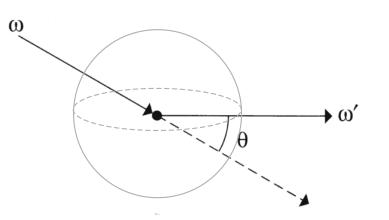
Isotropic

- Basic PFs
- PFs is constant



- Isotropic
 - Basic PFs
 - PFs is constant
 - Since
 - Area of sphere = 4?*r2
 - pfS are normalized (r =1)

$$p_{isotropic}(\mathbf{w} \rightarrow \mathbf{w'}) = \frac{1}{4\mathbf{p}}$$



Rayleigh

- Very small particles
- Acurately describes light scattering when
 - Particle radii < light wavelength</p>
- Good for atmospheric simulation

Mie

- Based on Maxwell's equations
- Broader range of particle sizes
- Good for fog and water droplets simulation

- Henyey and Greenstein
 - Easy to fit
 - Single control parameter
 - Controls relative proportion of forward backward scattering
 - $g \in (-1, 1)$
 - g < 0: back scattering</pre>

$$p_{HG}(\cos \boldsymbol{q}:g) = \frac{1}{4\boldsymbol{p}} \frac{1-g^2}{(1+g^2-2g(\cos \boldsymbol{q}))^{3/2}}$$

Increase complexity by combination

$$p(\cos \boldsymbol{q}) = \sum_{i=1}^{n} w_i p_{HG}(\cos \boldsymbol{q} : g_i)$$

Increase complexity by combination

$$p(\cos \boldsymbol{q}) = \sum_{i=1}^{n} w_i p_{HG}(\cos \boldsymbol{q} : g_i)$$

- More efficient version
 - Avoids 3/2 power computation

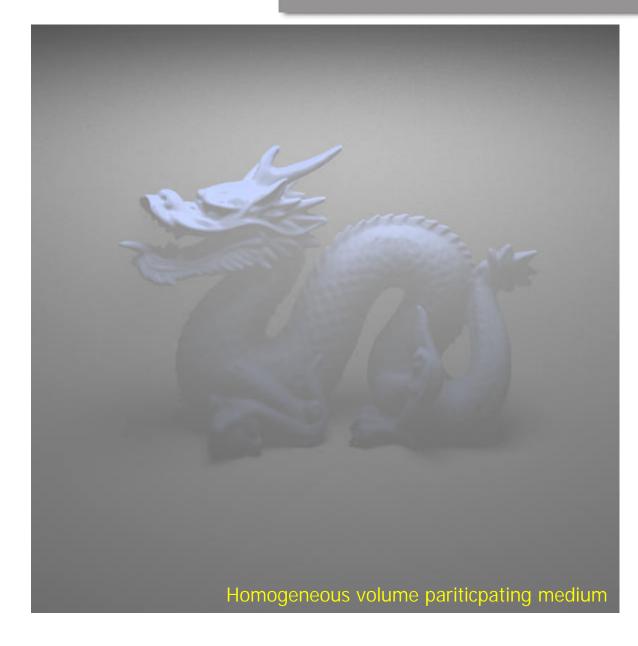
$$p_{Schlick}(\cos \boldsymbol{q}) = \frac{1}{4\boldsymbol{p}} \frac{1-k^2}{(1-k\cos \boldsymbol{q})^2}$$

Volume Interface and Homogeneous Media

- VolumeRegion class
 - Volume Scattering Interface in PBRT
 - General methods
 - BBox
 - IntersectP
 - Scattering information methods
 - Sigma_a (absorption)
 - Sigma_s (scattering)
 - Lve (emission)
 - P (phase value)
 - Sigma_t (attenuation)
 - Tau (optical thickness)

Volume Interface and Homogeneous Media

- Homogeneous volumes
 - Uniform particle density
 - One type of particle
- HomogeneousVolume : VolumeRegion
 - s_s And s_a are constant



- Inhomogeneous volumes
 - Still one type of particle
 - Spatially variable density
 - Scales scattering properties according to density, except for Tau
- DensityRegion : VoumeRegion
 - Method for obtaining density added
 - Tau not implemented
 - Dependent on shape and density distribution function

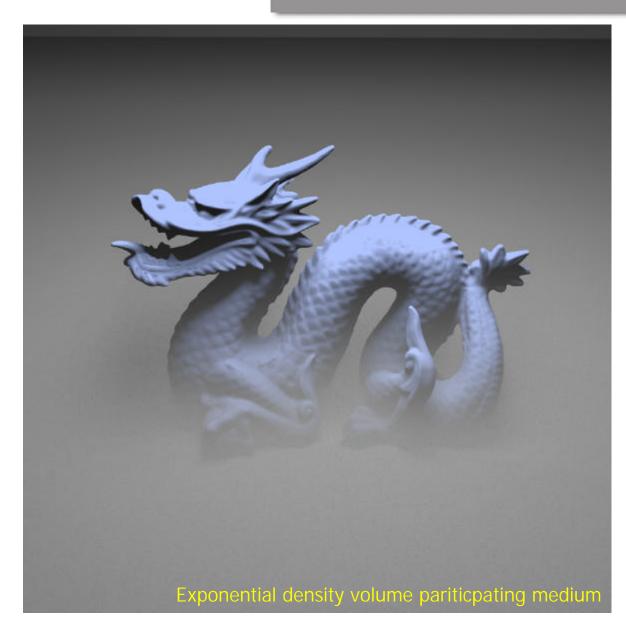
3D Grids

- VolumeGrid : DensityRegion
- Density values defined in a 3D matrix
- Intermediary values are interpolated
 - Manhatan distance
 - p1 (x1, y1) and p2 (x2, y2)
 Manhatan distance is |x1 x2| + |y1 y2|.
 - Trilinear interpolation
- Used in smoke pictures

- Exponential Density Volume
 - ExponentialDensity : Density Region
 - Density varies with height

$$d(h) = ae^{-bh}$$

Good for modelling Earth's atmosphere



Volume Aggregates

- AggregateVolume
 - Many Volume Regions
- Advantages
 - Simplifies scene
 - Easy integration with 3D structures
- Parameters creation
 - Sum of individual volumes` parameters

References

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- 5. Beer-Lambert law, Wikipedia
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